

method results, as can be seen in Table II. Comparisons at other values of atomic number and at other energies indicate that interpolated values of the parameters reproduce both dose rate and energy absorption build-up factors to within an average error of 5% and a maximum error of 20%.

## REFERENCES

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G. L. STROBEL\*

Bettis Atomic Power Laboratory†  
Pittsburgh, Pennsylvania  
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\* Present address: Douglas Aircraft Company, Inc., Missiles and Space Systems, 3000 Ocean Park Boulevard, Santa Monica, California.

† Operated for the U. S. Atomic Energy Commission by Westinghouse Electric Corporation.

### Re: "Thermal Neutron Flux Depression by Absorbing Foils" and "Flux Perturbations by Thermal Neutron Detectors"

Having received a number of useful comments about these two papers (1, 2), the authors would like to point out some corrections and limits of applicability of the two methods.

In the first paper, (1) the last line of paragraph 6 on page 301 should read . . . "identical with the  $\frac{3}{4}$  term in Eq. (7) as long as  $g \gg 1$ ." In Figs. 1, 2, and 3 the ordinates should

all be scaled up by a factor of 10, i.e., they should run from 7.6-8.6, 1.8-2.3, and 3.0-5.5 respectively. Equation (40) should read:

$$\frac{\bar{\phi}}{\phi_0} = \frac{[\frac{1}{2} - E_3(\tau)]/\tau}{1 + [\frac{1}{2} - E_3(\tau)] \cdot g_s \left( \frac{L}{a}, \frac{L}{\lambda} \right) \cdot \frac{g_v(L/\lambda, \tau)}{g_s(a = \infty, L/\lambda)}}$$

It should be noted that the variational method of the first paper should give exactly correct results for the cases of coins with very large radii. On the other hand, for zero radii coins Eq. (30) is zero although this does not give the expected flux ratio of 1.0 when inserted in Eq. (40). Thus for zero or small radii the equation above breaks down. However, this failure is due to the fact that the self-shielding factor  $[\frac{1}{2} - E_3(\tau)]/\tau$  is computed assuming that the foil radius  $a$  is much greater than its thickness  $t$ . One may avoid this difficulty by using the self-shielding factor computed by Skyrme for the case  $a \sim t$ . In that case, one would employ the expression  $\tau\{1 - \tau [A(g) - \frac{1}{2} \ln \sigma\tau]\}$  in place of  $[\frac{1}{2} - E_3(\tau)]$  in the equation above.  $A(g)$ , according to Skyrme, is given by

$$A(g) = \frac{3}{4} + 1/12 \pi g^3 + O(1/g^4)$$

for  $g = a \sum_{ad} \gg 1$

and

$$A(g) = 1/(\pi g) - \frac{1}{2} E_1(2g) + \frac{1}{2} (1 - \ln 2) + O(g)$$

for  $g \ll 1$ . Since these equations for  $A(g)$  were derived assuming  $\tau \ll 1$ , such an approximation has not the domain of validity possessed by the corrected form of Eq. (40).

The integral method of the second paper suffers from just the opposite difficulty. For very small radii coins the method works very well since integration over a very small volume of smooth functions proceeds with no difficulties. However, integration over large radii coins with a large number of radial points rapidly becomes very time-con-

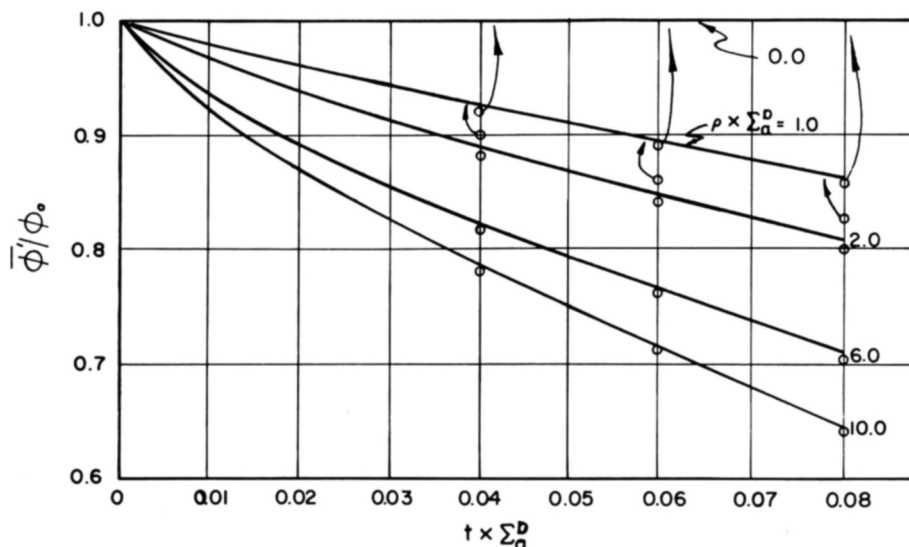


FIG. 1. The average normalized scalar flux in a coinshaped detector in water for various radii  $\rho$  and thicknesses as calculated by the variational and the integral methods.  $\circ$  variational method; — integral method.

suming and expensive. Thus it is not practical to extend the integral method to coins of very large radii.

The above comments should not be taken to indicate that there is little agreement between the two methods. Figure 1 shows that there is a large area of agreement between the two methods even when applied to detectors in water.

It should be noted that for the data in Fig. 1 scatter was assumed isotropic (i.e.  $\bar{\mu} = 0$ ) since most of the integral calculations were carried out on this basis. Calculations for a gold coin (of 5 mils thickness and 0.5 cm radius) in water to investigate the effect of anisotropy of scatter in water have been made using both methods. The results are as follows:

Average Scalar Flux in the Detector			
	$\bar{\mu}$	0.0	0.3
$\bar{\phi}/\phi_0$	(variational)	$0.773 \pm .005$	$0.803 \pm .005$
$\bar{\phi}/\phi_0$	(integral)	$0.813 \pm .002$	$0.831 \pm .002$

In view of the uncertainty in reading the graphs for use in the variational method and finiteness of the numerical integrations of the integral method, the agreement as to the sign and magnitude of the effect of anisotropy is quite encouraging. It is, however, unfortunate that the comparison was made in a region of small radii coins where the two methods do not agree too well in absolute magnitude.

Finally it should be emphasized that all the dimensionless plots in the second paper, i.e., Figs. 7-10 and 13-18, are not rigorously correct. They result from a compromise of about one percent between the dimensionless plots for gold and indium. Further investigation showed that the same dimensionless plots could be used for detector absorption cross section between 1.0 and 10.0 if one requires no more than plus or minus two or three percent accuracy. If, however, (1) high accuracy is required, (2) detector absorption is not considerably greater than its scatter, (3) scatter in the external medium is not isotropic, and (4) the detector is not in the size range covered, then one should return to the computer and calculate the particular cases of interest.

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Health Physics  
Oak Ridge National Laboratory  
Oak Ridge, Tennessee

Department of Nuclear Engineering  
University of Florida  
Gainesville, Florida

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R. H. RITCHIE

G. R. DALTON

### The Milne Problem with a Polynomial Source

The Milne problem with a source of the form  $x^n$  has been treated by a number of authors. Lundquist and Horak have expressed the emergent flux in terms of a recursion relation (1). Ueno has used the probabilistic approach to obtain the emergent flux in closed form (2). Busbridge has derived both the emergent flux and the angle integrated flux in the interior (3). The latter is obtained from an iteration procedure which is shown to converge to the correct solution. The purpose of this note is to derive closed expressions for the emergent angular distribution and the total flux by using a method described in ref. 4.

In plane geometry the energy independent transport equation for isotropic scattering in the laboratory system is

$$\mu \frac{d\psi}{dx} + \psi(x, \mu) = \frac{1}{2} \omega \int_{-1}^1 \psi(x, \mu') d\mu' + \frac{1}{2\Sigma} Q(x), \quad 0 \leq \omega < 1, \quad (1)$$

where  $\omega = \Sigma_s/\Sigma$ ,  $\mu$  is the direction cosine with the positive  $x$ -axis, and  $Q(x)$  is a volume source.  $x$  is measured in terms of the total mean free path. For a source of the form  $Q(x) = \exp sx$  the angular distribution  $\psi(0, -\mu)$ , ( $0 \leq \mu \leq 1$ ), of neutrons emerging from the surface  $x = 0$  of a semi-infinite slab can be shown to equal (2, 4)

$$\psi(0, -\mu) = (1/2\Sigma)H(-1/s)H(\mu)(1 - \mu s)^{-1}, \quad \mu \geq 0, \quad (2)$$

where  $H(\mu)$  satisfies the integral equation

$$H(\mu) = 1 + \frac{1}{2} \omega \mu H(\mu) \int_0^1 H(\mu')(\mu + \mu')^{-1} d\mu', \quad 0 \leq \mu \leq 1. \quad (3)$$

The  $H$ -functions have been discussed extensively (3, 5, 6). They are tabulated in the range  $0 < \omega \leq 1$  for small increments of  $\mu$  (5, 7). Their moments, defined as

$$h_n = \int_0^1 \mu^n H(\mu) d\mu, \quad (4)$$

are tabulated for  $0 \leq n \leq 20$  (7).

Expressing the source in terms of its Laplace transform,

$$Q(x) = \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} \bar{Q}(s)e^{sx} ds, \quad (5)$$

one finds for the angular distribution, according to Eq. (2),

$$\psi(0, -\mu) = \frac{1}{2\Sigma} H(\mu) \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} \bar{Q}(s)H\left(-\frac{1}{s}\right) \cdot (1 - \mu s)^{-1} ds, \quad \mu \geq 0. \quad (6)$$

where the contour must correspond to that of Eq. (5).

Equation (6) is quite general and can be applied to an arbitrary source  $Q(x)$ . Its use will be illustrated by applying it to a source of the form  $Q(x) = Qx^n$ . Inserting the Laplace transform of the source into Eq. (6) gives,

$$\psi^{(n)}(0, -\mu) = \frac{n!Q}{2\Sigma} H(\mu) \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} \frac{H(-1/s)}{s^{n+1}(1 - \mu s)} ds. \quad (7)$$

If the contour is closed by an arc of radius  $R$  in the left-hand plane, the contribution of the integral along the arc will vanish as  $R \rightarrow \infty$ .  $H(-1/s)$  and  $1/(1 - \mu s)$  are regular in the left-hand plane, so that the only singularities enclosed by the contour are at  $s = 0$ , where the integrand has a pole of